

СВОЙСТВА ПЛАЗМЫ С БОЗЕ КОНДЕНСАТОМ ЗАРЯЖЕННЫХ ЧАСТИЦ

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XII Марковские чтения

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Quite unusual results

found in our papers:

A.D., A. Lepidi, G. Piccinelli,

JCAP 0902 (2009) 027; Phys. Rev D, 80
(2009) 125009; JCAP 08 (2010) 031;

A.D., A. Lepidi Phys.Lett. A375(2011) 3188.

Similar results but by another method:

G. Gabadadze, R.A. Rosen,

Phys. Lett. B 658 (2008) 266;

JCAP 0810 (2008) 030;

JCAP 1004 (2010) 028.

Textbook formula for screening:

$$U(r) = \frac{Q}{4\pi r} \rightarrow \frac{Q \exp(-m_D r)}{4\pi r},$$

because the time-time component of the photon propagator acquires “mass”:

$$k^2 \rightarrow k^2 + \Pi_{00}(k) = k^2 + m_D^2,$$

where e.g. for relativistic fermions

$$m_D^2 = e^2 \left(T^2/3 + \mu^2/\pi^2 \right).$$

In presence of charged particle condensate the screening is not exponential but power law and oscillating as a function of distance.

We did not publish our work for about half a year, but then found out that an oscillating screening is known for plasma with degenerate fermions and is observed in experiment - Friedel oscillations.

Physics is different but qualitative behavior is the same.

Strangely until recently the effects on screening from condensate of a charged Bose field were not well studied, though it is a textbook problem.

Consider electrically neutral plasma with large electric charge density of fermions compensated by charged bosons. Bosons condense when their chemical potential reaches maximum value:

$$\mu_B = m_B.$$

Otherwise it is impossible to make larger asymmetry between bosons and antibosons.

Equilibrium distribution of condensed boson

$$f_B^C = C \delta^{(3)}(\mathbf{q}) + \frac{1}{\exp [(E - m_B)/T] \pm 1}$$

is a solution of the kinetic equation, it annihilates the collision integral for an arbitrary constant C .

f_{eq} is always determined by two parameters, either T and μ , or T and C , iff $\mu = m_B$.

Collision integral:

$$I_{coll} \sim |A_{fi}|^2 \Pi f_f \Pi(1 \pm f_i) - (\textit{inverse})$$

If T-invariance holds, i.e. $|A_{if}| = |A'_{fi}|$:

$$I_{coll} \sim [\Pi f_i (1 \pm f_f) - (i \leftrightarrow f)] d\tau.$$

$I_{coll} = 0$ for arbitrary T and C
iff $\mu = m$.

If T-invariance is broken and
 $|A_{if}| \neq |A'_{fi}|$, :

$$I_{coll}[f_{eq}] \sim \Pi f_i (1 \pm f_f) \left[|A_{fi}|^2 - |A_{if}|^2 \right]$$

This term is surely non-vanishing!

Do equilibrium distributions remain the same in T-broken theory?

Breaking of T-invariance is unobservable if only one reaction channel is open. In this case $T_{if} = T_{fi}^*$ with time reflected momenta.

f_B^C annihilates collision integral after summation over all relevant processes, due to S-matrix unitarity or CPT and conservation of probability.

Instead of the detailed balance condition there operates “the cyclic balance” condition

Screening properties of medium are expressed through f which is not necessarily equilibrium one. In calculations neither imaginary time method which may be inconvenient in presence of condensate or out of equilibrium, nor Matsubara-Keldysh technique are used. We started from the quantum equations of motion, solved them with Green's function up to e^2 order, and averaged corresponding operators not only over vacuum but also over “non-empty” medium.

Operator Maxwell equations:

$$\partial_\nu F^{\mu\nu}(x) = \mathcal{J}_B^\mu(x) + \mathcal{J}_F^\mu(x),$$

where bosonic current is

$$\mathcal{J}_B^\mu(x) = -ie[(\phi^\dagger(x)\partial^\mu\phi(x)) - (\partial^\mu\phi^\dagger(x))\phi(x)] + 2e^2 A^\mu(x)|\phi(x)|^2,$$

plus fermionic current:

$$\mathcal{J}_F^\mu(x) = e\bar{\psi}\gamma_\mu\psi.$$

Using equation of motion for quantum operator ϕ :

$$(\partial^2 + m^2)\phi(x) = \mathcal{J}_\phi(x)$$

express ϕ through A_μ :

$$\phi(x) = \phi_0(x) + \int d^4y G_B(x-y)\mathcal{J}_\phi(y),$$

ϕ_0 is free field operator. In the lowest order in e take $\phi = \phi_0$ in $\mathcal{J}_B^\mu(x)$.

The r.h.s. of the Maxwell equations in e^2 order is linear (but non-local) in A_μ and bilinear in ϕ_0 and ψ_0 .

Expand free fields as usually:

$$\phi_0(x) = \int d\tilde{q} \left[a(q) e^{-iqx} + b^\dagger(q) e^{iqx} \right].$$

Average over medium:

$$\begin{aligned} \langle a^\dagger(q) a(q') \rangle &= f_B(E_q) \delta^{(3)}(q - q'), \\ \langle a(q) a^\dagger(q') \rangle &= [1 + f_B(E_p)] \delta^{(3)}(q - q'). \end{aligned}$$

Unity is subtracted, since it is vacuum contribution.

The Fourier transform of the Maxwell equations in plasma is:

$$\left[k^2 g^{\mu\nu} - k^\mu k^\nu + \Pi^{\mu\nu}(k) \right] A_\nu(k) = \mathcal{J}^\mu(k)$$

where the boson contribution is:

$$\Pi_{\mu\nu}^B(k) = e^2 \int \frac{d^3q}{2(2\pi)^3 E} [f_B(E, \mu) + \bar{f}_B(E, \bar{\mu})] \left[\frac{l_\mu l_\nu}{l^2 - m^2} + \frac{p_\mu p_\nu}{(p^2 - m^2)} - 2g_{\mu\nu} \right]$$

where $l = k + q$, $p = k - q$, and $E = \sqrt{q^2 + m^2}$.

Solving Fourier transformed the linear Maxwell equation for A_t :

$$\Pi_{tt}(0, k) = \frac{e^2}{2\pi^2} \int_0^\infty \frac{dq q^2}{E_B} [f_B(E_B, \mu_B) + \bar{f}_B(E_B, \bar{\mu}_B)] \left[1 + \frac{E_B^2}{kq} \ln \left| \frac{2q + k}{2q - k} \right| \right],$$

plus similar contribution from fermions which neutralize the plasma.

This is the well known result for Π_{tt} in order e^2 .

The screened Coulomb potential is the Fourier transform of tt-component of the photon Green's function in medium:

$$U(r) = e^2 \int \frac{d^3k}{(2\pi)^3} \frac{e^{i\mathbf{k}\cdot\mathbf{r}}}{k^2 + \Pi_{tt}(k)} = \frac{e^2}{2\pi^2 r} \int_0^\infty dk k \frac{\sin kr}{k^2 + \Pi_{tt}}.$$

Asymptotics of the potential of charged impurities is determined by the singularities of Π_{tt} in complex k -plane.

Comment.

Singularities of $f(z)$:

$$f(z) = \int_a^b dy g(z, y)$$

in complex z -plane appear at such z for which singularities of $g(z, y)$, i.e. $y_c(z)$, in complex y -plane coincides with the bounds of integration, a or b , or $y_c(z)$ pinches the contour of integration.

Two types of singularities:

1. Poles of $[k^2 + \Pi_{tt}(k)]^{-1}$. E.g. Debye pole. Necessary to check that the position of the poles are at small k , such that the infrared asymptotics of Π_{tt} is valid.
2. Singularities of $\Pi_{tt}(k)$, originating from the pinch of the integration contour in q -plane by poles of f and by branch points of \log .

Without condensate one obtains the usual k -independent Debye screening:

$$\Pi_{tt}(0, k) = m_D^2$$

originating from a pole at imaginary axis of k .

With condensate the corrections to Π_{tt} at low k are infrared singular:

$$\frac{\Delta\Pi_{tt}}{e^2} = \frac{m_B^2 T}{2k} + \frac{C}{(2\pi)^3 m_B} \left(1 + \frac{4m_B^2}{k^2} \right)$$

Both terms in the r.h.s. appear only if $\mu = m_B$.

Instead of exponential the screening becomes power law and oscillating, depending upon parameters, m_j :

$$\Pi_{tt} = m_0^2 + m_1^3/k + m_2^4/k^2.$$

May this have something to do with confinement
Recent paper: P. Gaete, E. Spalucci, 0902.0000
– confinement in Higgs phase.

Contribution from poles in the limit of large $m_2 r$ but when power law terms are subdominant:

$$U(r)_{pole} = \frac{Q}{4\pi r} \exp(-\sqrt{e/2m_2}r) \times \cos(\sqrt{e/2m_2}r).$$

Oscillating screening is known for **degenerate** fermions - Friedel oscillations. Observed in experiment.

The screening electrons are waves with $k = k_F$ (from B. Shklovsky).

Comment.

Friedel oscillations are commonly believed to be zero T phenomenon, because in this case the integral over q is in finite interval and the singularity in k appears when log branch point coincides with the upper limit of the integration.

However the "pinch" method works at $T \neq 0$ and the $T = 0$ limit can be recovered by summing all the singularities. Non-zero T corrections, absent in textbooks can be obtained in this way.

Contribution from the integral along imaginary axis is nonzero because Π_{00} contains an odd in k term. If $m_2 \neq 0$, the dominant term is

$$U(r) = -\frac{12Qm_1^3}{\pi^2 e^2 r^6 m_2^8}.$$

If $T \neq 0$, $\mu = m_B$, but the condensate is not yet formed, the asymptotic decrease of the potential becomes:

$$U(r) = -\frac{Q}{\pi^2 e^2 r^4 m_1^3} = -\frac{2Q}{\pi^2 e^2 r^4 m_B^2 T}.$$

Contribution from logarithmic cuts (analogous to Friedel oscillations for fermions).

If the first “pinch” (between the poles of $f(q)$ and logarithmic branch point) dominates:

$$U_1(r) = -\frac{32\pi Q}{e^2 m_B r^2} \frac{e^{-z}}{\ln^2(2\sqrt{2}z)} \sin z ,$$

where $z = 2r\sqrt{2\pi T m_B}$.

NB: $U_1(r)$ is inversely proportional to e^2 and formally vanishes at $T \rightarrow 0$, but remains finite if $\sqrt{T m_B} r \neq 0$.

All pinches are comparable:

$$U(r) \approx -\frac{3Q}{2e^2 T^2 m_B^3 r^6 \ln^3(\sqrt{8m_B T} r)}.$$

$U \sim T^{-2}$ valid if $r \ll 1/\sqrt{16\pi T m_B}$,
i.e. if $T = 0.1\text{K}$ and $m_B = 1\text{GeV}$ the
distance should be bounded from above
as $r \ll 3 \cdot 10^{-8}$ cm.

Condensation of vector bosons.

W^\pm would condense in the early universe if lepton asymmetry was sufficiently high.

It leads to large electric asymmetry of W , such that $\mu_W = m_W$.

Plasma neutrality was maintained by quarks and leptons.

Vector bosons have additional degrees of freedom, their spin states, and their condensation demonstrates richer possibilities: Depending on the sign of the pairwise spin-spin coupling W 's would condense either in $S = 0$ (scalar) state or in $S = 2$ (ferromagnetic) state.

Magnetic spin-spin interaction through one photon exchange (similar to Breit equation):

$$U_{em}^{spin}(r) = \frac{e^2 \rho^2}{4\pi m_W^2} \left[\frac{(S_1 \cdot S_2)}{r^3} - \frac{3}{3} \frac{(S_1 \cdot r)(S_2 \cdot r)}{r^5} - \frac{8\pi}{3} (S_1 \cdot S_2) \delta^{(3)}(r) \right].$$

Here ρ is the ratio of magnetic moment of W to the standard one.

For S -wave the energy is shifted by the last term only.

Local quartic self-coupling of W :

$$U_{4W}^{(spin)} = \frac{e^2}{8m_W^2 \sin^2 \theta_W} (S_1 S_2) \delta^{(3)}(r).$$

The net result $U_{em} + U_{4W}$ is negative, so $S = 2$ state is energetically favorable and spontaneous magnetization in the early universe is possible.

Suppression due to screening.

The ij component of W propagator probably remains massless: $\Pi_{ij} \sim 1/q^2$. In QED it is true in perturbation theory, while in non-Abelian theories the screening may occur in higher orders of perturbation theory due to infrared singularities. The screening would diminish the long-range ferromagnetic spin-spin coupling while the local W^4 coupling is not screened.

If the propagator is modified, and the wave function of W -bosons is constant in space, the spin-spin energy shift is:

$$\delta E \sim \int \frac{d^3 q \delta(q)}{(2\pi)^3} \frac{q^2 (S_1 S_2) - (q S_1)(q S_2)}{q^2 + \Pi_{ss}(q)}$$

$\delta E = 0$, if $\Pi_{ss} \neq 0$ at $q = 0$.

However, the integration over space should be done with an upper limit, l , equal to the average distance between the W bosons so instead of $\delta^{(3)}(q)$, we obtain:

$$\int_0^l d^3r e^{iqr} = \frac{4\pi}{q^3} [\sin(ql) - ql \cos(ql)].$$

and the energy shift is non-zero:

$$\delta E = -\frac{(S_1 S_2) e^2}{l^3 m_W^2} F(l),$$

$$F(l) = \int_0^\infty \frac{dx [x \sin x + l^2 \Pi_{SS} \cos x]}{x^2 + l^2 \Pi_{SS}(x/l)}.$$

If $l^2\Pi_{ss}$ is nonnegligible the e.m. part of the spin-spin interaction would be suppressed and the ferromagnet turns into an antiferromagnet. This might happen at T above the EW phase transition when the Higgs condensate is destroyed and $m_{W,Z}$ appear as a result of temperature and density corrections and are relatively small.

The quantitative statement depends upon the **(unknown)** modification of the space-space part of the photon propagator in presence of the Bose condensate of charged W – **a problem to solve.**

Problem of large scale magnetic fields:
 $B \sim \mu G$ at several kpc. In the intergalactic space the fields are probably 2-3 orders of magnitude weaker, but still non-vanishing
Dynamo operates only in galaxies.
Maybe ferromagnetism of W might create seeds for large scale magnetic fields.

Screening of magnetic fields is connected with the space-space components, which, in the homogeneous and isotropic case is

$$\Pi_{ij} = a(k) \left(\delta_{ij} - \frac{k_i k_j}{k^2} \right) + b(k) \frac{k_i k_j}{k^2}.$$

Multiplying Π_{ij} by δ_{ij} and by $k_i k_j$ we obtain $b(k) = 0$ and

$$a(k) = \frac{e^2}{32\pi^3} \int \frac{d^3q}{E} (f + \bar{f}) \left[2 + \frac{2k^2(4q^2 - k^2)}{4(\tilde{k}\tilde{q})^2 - k^4} \right].$$

where $\tilde{k}\tilde{q} = kq \cos \theta$.

If only the condensate term is retained:

$$a^{(C)}(k) = \frac{e^2 C}{8\pi^3 m_B} \equiv e^2 m_C^2.$$

Since $a(0) = \text{const} \neq 0$, the magnetic field is exponentially screened. In absence of magnetic monopoles magnetic field can be screened only by currents, **hence plasma with BEC of electrically charged Bose field must be superconductive-** well known result. (Two regimes of superconductivity: weakly coupled Cooper pairs, i.e. BCS or strong coupling BEC regime.)

If $\mu < m_B$, then $\Pi_{ij}(k)$ vanishes as k^2 in the limit $k \rightarrow 0$, as expected:

$$a(k) \approx \frac{e^2 k^2}{24\pi^2} \int \frac{dq}{E} (f + \bar{f})$$

and magnetic fields are not screened.

If $\mu = m_B$, even without condensate, i.e. at $C = 0$, $a(k)$ vanishes only as a first power of k , which leads to unusual screening features.

$a(k)$ is singular in the limit $m_B = 0$, since the integral diverges as $1/q^2$ at the lower limit of integration, $q = 0$. Moreover, a singularity at $k = 0$ exists for massive particles if $\mu = m_B$. The singularity comes from the integration region where $q \sim k$ due to singularity of $f(q)$ at low q . So we obtain for $k \rightarrow 0$:

$$a^{(sing)}(k) = \frac{e^2 T}{16} k.$$

For small k this term would dominate over the usual k^2 term and change the screening behavior.

In the transverse gauge, $k_j A_j = 0$, the Maxwell equation can be solved as

$$A_i(x) = \int d^3y G(x - y) \mathcal{J}_i(y).$$

The asymptotics of $G(r)$ at large r is determined by

$$G(r) = \frac{(-i)}{4\pi^2 r} \int_0^\infty dk k \frac{(e^{ikr} - e^{-ikr})}{k^2 + a(k)}.$$

$a(k)$ may contain odd terms in k , so the integral along the half real k -axis cannot be extended to the whole real axis.

It leads to non-canonical screening terms.

Since $a(k) = k^2 + e^2 m_C^2 + e^2 T k / 16$, the integral can be rewritten as:

$$G(r) = \frac{(-i)}{4\pi^2 r} \int_0^\infty dk k \left(e^{ikr} - e^{-ikr} \right) \frac{(k^2 + e^2 m_C^2 - e^2 T k / 16)}{(k^2 + e^2 m_C^2)^2 - e^4 T^2 k^2 / 256}.$$

The integral of the even part is expressed through the residues of the poles in the complex k -plane at:

$$k^{(pole)} = \pm i \sqrt{e^2 m_C^2 - \frac{e^4 T^2}{1024}} \pm \frac{e^2 T}{32}.$$

If $m_C > e^2 T / 32$, the screened potential would be exponentially cut with superimposed oscillations. For $e^2 T \ll 32 m_C$, the Green function takes the form:

$$G(r) \sim \exp(-em_C r) \cos(e^2 r T / 32).$$

In this case the spatial damping scale is much shorter than the oscillation scale. However, for $eT \sim m_C$ the scales are comparable.

The contour of the integration of odd in k , part can be closed in upper or lower quadrant of the complex k -plane. So in addition to the poles in these quadrants the contributions from the integrals over the imaginary axis are to be included. **They produce a power law screening.**

If $C = 0$, but $\mu = m_B$, then at small k : $a(k) \approx e^2 k T / 16$, so the Green's function drops as: **$G(r) \sim 8 / (\pi^2 e^2 r^2 T)$** . This is realized when $r > 1/T$.

In presence of condensate the Green's function acquires an additional constant term $e^2 m_C^2$. In this case the contribution of the integral over the imaginary axis of k gives $G \sim T / (16 e^2 \pi^2 r^4 m_C^4)$.

Changing of the asymptotics of screening signals formation of the condensate.

THE END

Calculation of singularity.

It is convenient to separate the integral into two parts $0 < q < k/2$ and $k/2 < q < \infty$. In the first part we introduce the new integration variable $x = 2q/k$, so $0 < x < 1$. In the limit of small k the energy can be expanded as $E_B \approx m_B + k^2 x^2 / 8m_B$. At small q the distribution function is infrared singular:

$$\left[\exp \left(\frac{E_B - m_B}{T} \right) - 1 \right]^{-1} \approx \frac{2m_B T}{q^2} = \frac{8m_B T}{k^2 x^2}.$$

Usually this singularity is not dangerous because it is canceled by the integration measure, $\sim q^2$. However, the logarithmic term behaves as k/q for $q > k$ and as q/k for $q < k$. Thus the integral is finite, but it does not vanish as k^2 when $k \rightarrow 0$.

The first part of the integral with $q < k/2$ can be taken analytically and we obtain:

$$a_1^{(s)}(k) = \frac{e^2 k T}{8\pi^2} \int_0^1 dx \left[2 - \left(x - \frac{1}{x} \right) \ln \left| \frac{1+x}{1-x} \right| \right]$$

$$\frac{e^2 k T}{8\pi^2} \left(1 + \frac{\pi^2}{4} \right).$$

There is also another contribution coming from the part of the integral with $q > k/2$. As $k \rightarrow 0$, the second part of the integral, $k/2 < q < \infty$, gives:

$$a_2^{(s)}(k) = \frac{e^2 k T}{8\pi^2} \int_1^\infty dx \left[2 - \left(x - \frac{1}{x} \right) \ln \left| \frac{1+x}{1-x} \right| \right] \\ \frac{e^2 k T}{8\pi^2} \left(-1 + \frac{\pi^2}{4} \right),$$

such that the total contribution is:

$$a^{(s)}(k) = a_1^{(s)}(k) + a_2^{(s)}(k) = \frac{e^2 T}{16} k. \quad (4)$$

For small k this term could dominate over the usual k^2 term and would change the screening behavior.

so we present the denominator as half of sum and difference of even and odd function as following:

$$f(k) = [f(k) + f(-k)]/2 + [f(k) - f(-k)]/2 \quad (5)$$

Since $a(k) = k^2 + e^2 m_C^2 + e^2 T k / 16$, eq. (??) can be rewritten as:

$$G(r) = \frac{(-i)}{4\pi^2 r} \int_0^\infty dk k \frac{\left(e^{ikr} - e^{-ikr} \right) \left(k^2 + e^2 m_C^2 - e^2 T k / 16 \right)}{\left(k^2 + e^2 m_C^2 \right)^2 - e^4 T^2 k^2 / 256}$$

The integral of the even part may be transformed, as usually, into the integral along the whole real axis and after closing the contour in the upper (for e^{ikr}) or lower (for e^{-ikr}) half-plane we express the result through the residues in the corresponding poles in the complex k -plane at:

$$k^{(pole)} = \pm i \sqrt{e^2 m_C^2 - \frac{e^4 T^2}{1024}} \pm \frac{e^2 T}{32}. \quad (7)$$

If $m_C > e^2 T/32$, the resulting screened potential would be exponentially cut with superimposed oscillations. For $e^2 T \ll 32 m_C$, the Green function takes the form:

$$G(r) \sim \exp(-em_C r) \cos(e^2 r T/32). \quad (8)$$

In this case the spatial damping scale is much shorter than the oscillation scale. However, if $eT \sim m_C$ the scales are comparable.